

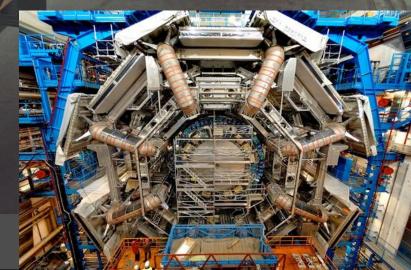
Searches for Supersymmetry in ATLAS

Pawel Brückman de Renstrom Institute of Nuclear Physics P.A.N. on behalf of the **ATLAS** Collaboration

Epiphany Conference

on Physics in Underground Laboratories and its Connection with LHC Kraków,07/01/2010





MOTIVATION

see Yann Mambrini's talk for more details

Boson-fermion unification appears "sexy" to some, but this is a matter of taste*. There are indisputable facts, however:

SUSY provides remedies for large number of fundamental problems of to-date physics (hierarchy problem, m_{Higgs} fine tuning, running of coupling const.,

connection to the String Theory, DM candidate), being at the same time highly predictive (at least in its minimal incarnation).

The most relevant aspect in the context of this conference is admittedly the existence of the DM candidate thanks to the R-parity conservation. This itself puts constraints on models and parameter space itself. The LSP must be:

weakly interacting,

➤ neutral.

The ONLY problem there is with SUSY is that it must be a broken symmetry and nobody has ever seen any direct evidence for superpartners The Supersymmetry breaking mechanism has implications on the low-E scale physics:

mSUGRA appears among the most popular scenarios for Supersymmetry breaking with lightest neutralino LSP (in the great majority of parameter space).
 a hope for direct WIMP detection a for Supersymmetry breaking alternative with gravitino LSP and lightest neutralino or a sfermion (usually stau) NLSP.

Other scenarios as split SUSY (R-hadrons), AMSB, NUHM are also considered in ATLAS but will not be covered here.
*as in life



Results shown here obtained if the framework of the final assessment of ATLAS physics potential and reported in the CERN-OPEN-2008-020 are normalised to 1fb⁻¹ at the CM energy of 14 TeV.

Results assume conservative uncertainties on the SM backgrounds (20% for top and electroweak, 50% for QCD).

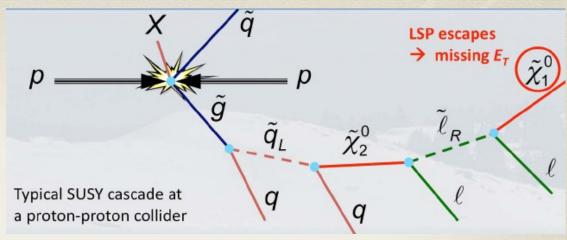
Scaling with luminosity is difficult due to importance of datadriven background estimation which will also depend on statistics – ongoing effort.



Generic SUSY signatures at a hadron collider

R-parity conserving SUSY:
pair-produced
two LSP's in the final state
if WIMP-like must be
neutral and weakly
interacting => escapes
detection!

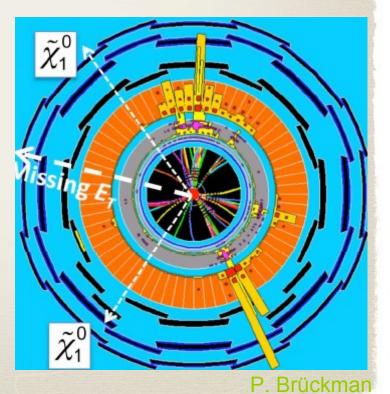
N.B. not unique to SUSY (UED, Little Higgs, ...)

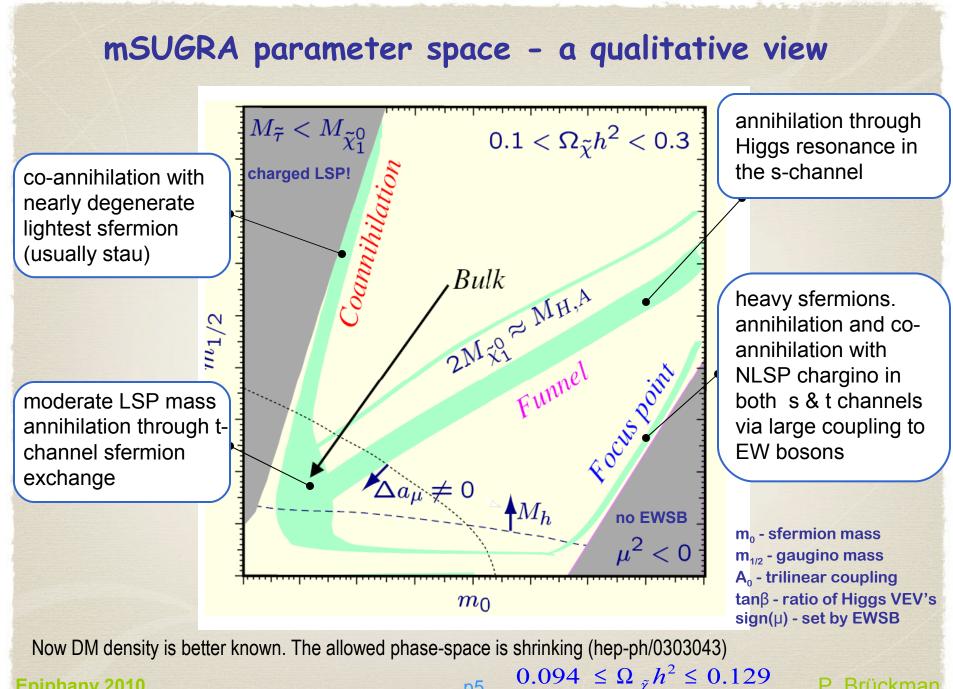


Inherent difficulty:

No direct mass resonance sparticles
 Preferably strongly produced => final states complicated (multi-jet + leptons + ∉_T)
 This implies large MC uncertainties on the backgrounds (mostly ttbar, W&Z+jets, QCD)
 Data-driven methods necessary!
 Rescue from:

- High transverse energy (M_{eff}) in the event
- ✓ Isolated leptons





Epiphany 2010

p5

ATLAS choice of mSUGRA benchmark points

$M_{1/2}$ [GeV A_0 [GeV $\sigma_{\rm LO}$ [pb] M_0 [GeV $\tan \beta$ $\arg \mu$ SU1 $\overline{70}$ 3500 10 8.15+Br(20 → (1) > 0.15 SU2 35503000 105.17+10020.851400 1400 SU3 300-3006 +m_n = 122 GeV τ̃⁰ LSP SU4 200160-40010+294.46SU6 320 3750 50+4.47Olu SU8.1 402103600 +6.481200 1200 SU1 SU2 SU3 SU4 SU6 SU8.1 SU9 Particle 419.84 764.90 3564.13 636.27 870.79 801.16 956.07 d_L 1000 1000 3563.24 631.51 412.25 866.84 797.09 952.47 \tilde{u}_L 760.42 m_{1/2} (GeV) \tilde{b}_1 697.90 2924.80 575.23 358.49 716.83 690.31 868.06 2131.11 424.12 206.04 641.61 603.65 725.03 Ťı 572.96 đ₽ 733.53 3576.13 610.69 406.22 840.21 771.91 920.83 $m_{\rm h} = 120 \, {\rm GeV}$ 800 3574.18 404.92 842.16 923.49 800 \tilde{u}_R 735.41 611.81 773.69 3500.55 399.18 779.42 910.76 Ъэ 722.87 610.73 743.09 SU8.1 \tilde{t}_2 749.46 2935.36 650.50 445.00 797.99 766.21 911.20 $Br(\tilde{\chi}^0_{s}\rightarrow h^0\tilde{\chi}^0_{s}) > 0.5$ 255.13 3547.50 230.45 231.94 411.89 325.44 417.21 \tilde{e}_L Ñe. 238.31 3546.32 216.96 217.92 401.89 315.29 407.91 600 600 malemal 146.50 3519.62 149.99 200.50 181.31 151.90 320.22 $\tilde{\tau}_1$ SU9 3532.27 215.53 358.26 296.98 401.08 \tilde{v}_{τ} 237.56216.29 154.06 3547.46 155.45 212.88 351.10 253.35 340.86 \tilde{e}_R 256.98 3533.69 232.17 236.04 392.58 331.34 416.43 $\tilde{\tau}_2$ 400 SU₂ 400 ĝ 832.33 856.59 717.46 413.37 894.70 856.45 999.30 O SU6 $\tilde{\chi}^{0}_{1} \tilde{\chi}^{0}_{2} \tilde{\chi}^{0}_{3} \tilde{\chi}^{0}_{4}$ $\tilde{\chi}^{0}_{3} \tilde{\chi}^{0}_{4}$ 136.98 103.35 117.91 59.84 149.57 142.45 173.31 SU₃ $Br(\gamma_0^0 \rightarrow Z^0 \gamma_1^0) > 0.5$ 325.39 263.64 160.37 218.60 113.48 287.97 273.95 179.76 466.44 463.99 308.94 477.23 463.55 520.62 m_b = 114 GeV 200 **SU4** 200 m = 103 GeV 294.90 327.76 492.23 479.01 536.89 483.30 480.59 $\tilde{\chi}_1^+$ 149.42 262.06 218.33 113.22 288.29 274.30 326.00 $\tilde{\chi}_{2}^{+}$ 326.59 479.22 536.81 483.62 286.81 480.16 492.42 ron NO EWSB h^0 115.81 119.01 114.83 113.98 116.85 116.69 114.45 H^0 370.47 0 515.99 3529.74 512.86 388.92 430.49 632.77 1200 0 200 400 600 800 1000 1400 1600 1800 2000 A^0 512.39 3506.62 511.53 368.18 386.47 427.74 628.60 H^+ 521.90 3530.61 518.15 378.90 401.15 440.23 638.88 m_o (GeV) 175.00 175.00 175.00 175.00 175.00 175.00 175.00 t

Exclusion regions for $A_0 = 0$, $\mu > 0$, tan $\beta = 10$

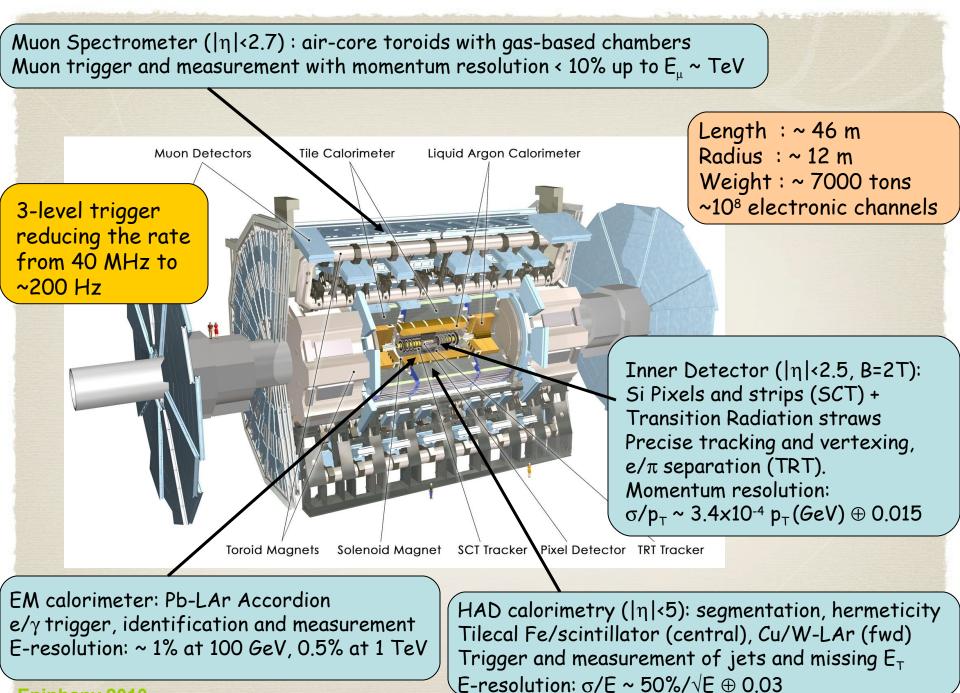
Epiphany 2010

Mt Blanc

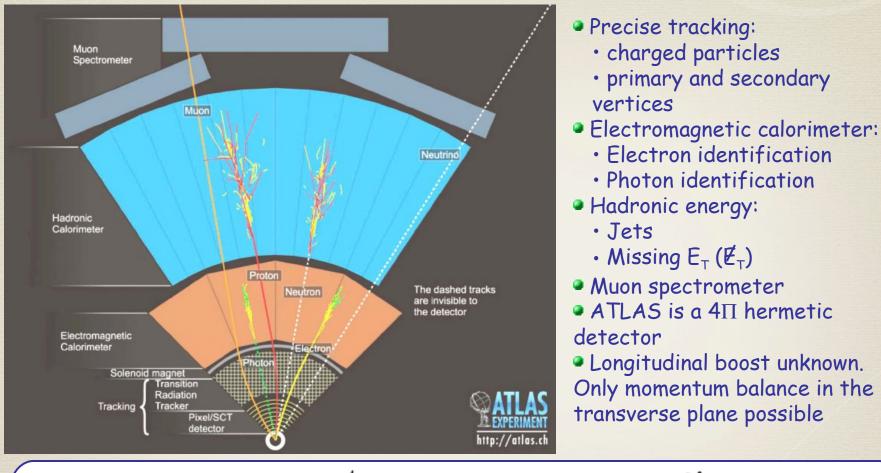


• pp collision cm : 14 TeV (x7 Tevatron)

25 ns bunch spacing
1.1 10¹¹ proton/bunch
Design lumi: 10³⁴cm⁻²s⁻¹ (10 nb⁻¹s⁻¹)
Physics/year ≈ 100 days
100 fb⁻¹ /year; ≈20 int./x-ing (≥2012?)
Low lumi: <10³⁵cm⁻²s⁻¹ (1 nb⁻¹s⁻¹)
10 fb⁻¹ /year ; ≈2 int./x-ing (≥2011)
Initial lumi:<10³²cm⁻²s⁻¹ (0.1 nb⁻¹s⁻¹) (2010)
<1 fb⁻¹ /year (at reduced energy!)



Measurements from the ATLAS spectrometer



Definition of some variables relevant to the discussed analyses

$$M_{\text{eff}} \equiv \sum_{i=1}^{4} p_T^{\text{jet},i} + \sum_{i=1} p_T^{\text{lep},i} + E_T^{\text{miss}} \qquad S_T \equiv \frac{2\lambda_2}{(\lambda_1 + \lambda_2)} S_{ij} = \sum_k p_{ki} p^{kj}$$

 $M_T^2(\mathbf{p}_T^{\alpha}, \mathbf{p}_T^{\text{miss}}, m_{\alpha}, m_{\chi}) \equiv m_{\alpha}^2 + m_{\chi}^2 + 2\left(E_T^{\alpha}E_T^{\text{miss}} - \mathbf{p}_T^{\alpha} \cdot \mathbf{p}_T^{\text{miss}}\right)$

mSUGRA – generic searches in ATLAS:

4 jets + 0 leptons 4 jets + 1 lepton 2 (or 3) jets + 0 leptons 2 (or 3) jets + 1 lepton o 4 jets + 2 leptons (OS or SS) o 3 leptons + jet ✓ 3 leptons + ∉_T **ο** 4 jets + τ • 4 jets + \geq 2 b-tagged

Enhanced at $tan\beta >>1$

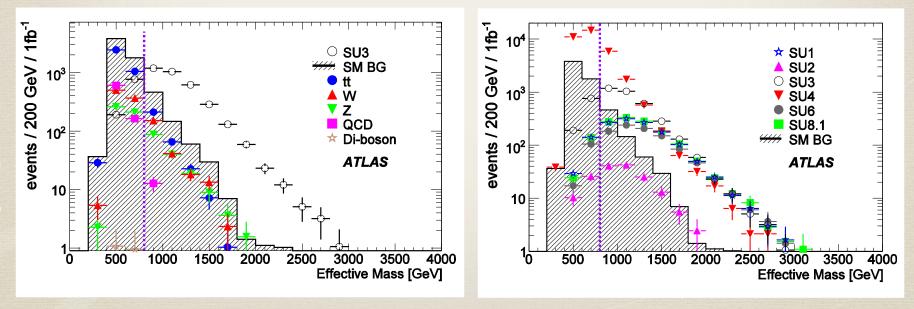
P. Brückman

The search strategy is largely <u>motivated</u> by the cosmological constraints on the DM relic density. Nevertheless, the final sensitivity scan is generic and takes into account only the direct exclusion limits from LEP and Tevatron. DM constraints merely a hint rather than constraint to generic SUSY searches!

mSUGRA: inclusive O-lepton mode

This channel has the highest statistics but suffers from remaining QCD background
 Cuts on sphericity & colinearity of ∉_T with a jet reduce QCD contribution.

- ttbar background remains dominant after all cuts.
- Only SU2 (focus point) cannot be assessed in this channel. Need leptonic signatures!

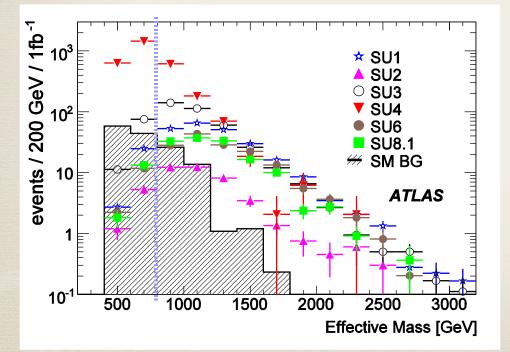


mSUGRA: inclusive 1-lepton mode

 QCD background efficiently suppressed by the isolated lepton requirement.
 Itbar background dominates by far after all cuts.

Similar significance to the 0-lepton mode.

□ SU2 (focus point) more accessible.



1 isolated lepton p_T>20GeV no other leptons (p_{τ} >10GeV) ≥4 jets p₇>50GeV ≥1 jet p₋>100GeV <mark>∉_⊤>100GeV</mark> ₽_>0.2M_{eff} S_⊤>0.2 $\Delta \phi(\text{jet}_{1,2,3}-\not\!\!E_T)>0.2$ **M_τ>100GeV** M_{eff}>800GeV trigger: J70_xE70

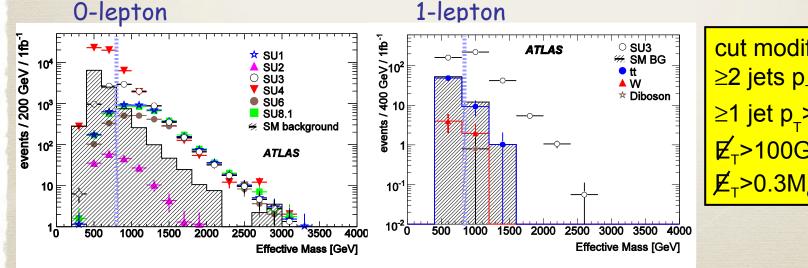
P. Brückman

mSUGRA: lower jet multiplicity option

Supersymmetry generically includes lower jet multiplicity final states, e.g.:

□ Lower jet multiplicity may be attractive especially for the early data where understanding of topologically complicated events may be limited (systematic uncertainties in modelling of multi-parton final states).

□ The approach proves efficient in both 0 and 1 lepton modes.

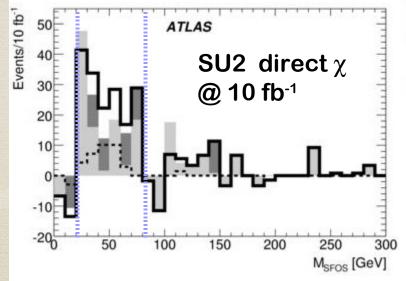


piphany 2010

cut modification: $\geq 2 \text{ jets } p_{T} > 100 \text{GeV}$ $\geq 1 \text{ jet } p_{T} > 150 \text{GeV}$ $\not{ E}_{T} > 100 \text{GeV}$ $\not{ E}_{T} > 0.3 \text{M}_{eff}$

mSUGRA: tri-lepton analysis (direct gaugino production)

- □ "Worst case scenario": very high m₀ and suppressed strong production through gluinos.
- Assessed using SU2 benchmark point with a jet veto.
- Direct chargino-neutralino production with three leptons in the final state has small cross-section (32.6 fb).
- SM background low but very stringent lepton isolation cuts required to suppress ttbar and Zb. WZ is inherently irreducible. Systematics small (5%).
 Requires considerably larger L_{int.} (5σ discovery with ~80 fb⁻¹)



pair of OSSF leptons (e or μ) NO 81.2<m_{OSSF}<102.2 GeV \geq 3 leptons pT>10 GeV $p_{T \text{ track,max}}^{\Lambda R=0.2}$ >2 Gev for electron $p_{T \text{ track,max}}^{\Lambda R=0.2}$ >1 Gev for muon \not{E}_{T} >30 GeV no jet with p_T>20GeV trigger: L2_e22i || L2_mu20

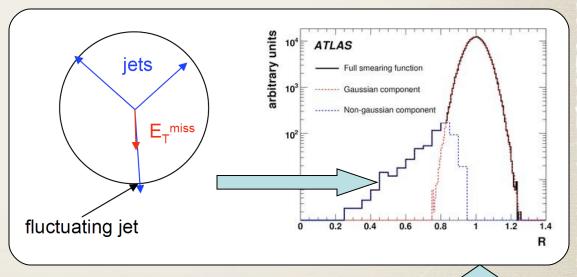
Name of the game: Understanding SM backgrounds from real data!

SM backgrounds must eventually be estimated from data itself. Strategy includes:

➤ excellent understanding of the detector response (lepton efficiencies, JES, ∠_T...)
 ➤ understanding of the individual backgrounds using control samples
 Different techniques used for different types of backgrounds and analysis modes.

Example: jet response

The non-Gaussian tail of the resolution function obtained from the jet mismeasurements in the "mercedes" QCD events



P. Brückmai

The Gaussian part of the resolution function from the jet-photon balance

Data-driven methods - the guiding principle

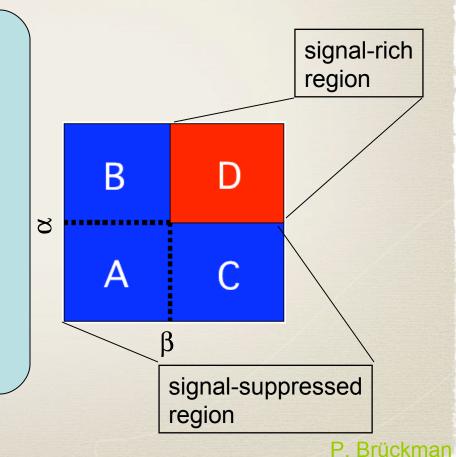
Most have a common principle which relies on identifying a control region (exclusive to the actual signal search region) which is signal suppressed but still representative for the background.
 One needs two variables (α,β) which are approximately independent:

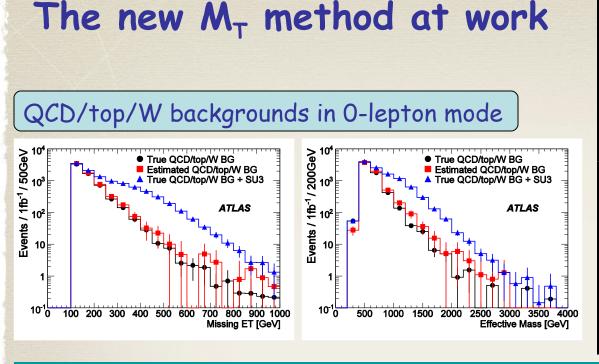
> β distribution from background events in the signal region can be estimated from data using:

≻D=CxB/A

normalisation from data-suppressed region using an independent α variable

If there is no SUSY signal (signal region consistent with the predicted background) -> DONE
 > Otherwise one can iterate subtracting the observed signal from the control sample ("new M_T method")



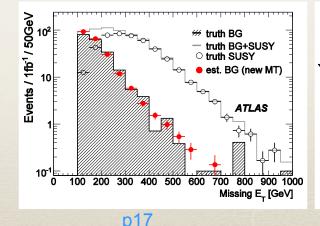


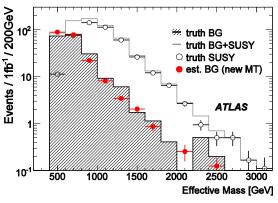
ttbar & W+jets: require a lepton (p_T >20GeV) drop it and recalculate kinematics normalise in the region: 100 GeV< $\not E_T$ <200 GeV

QCD:

top/W backgrounds in 1-lepton mode

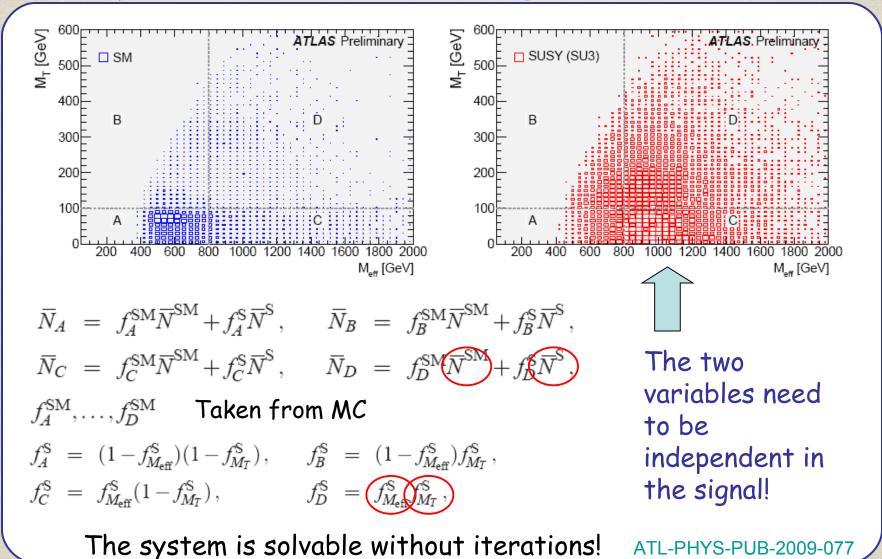
 M_{T} (lepton, $\not{\!\!E}_{T}$) control: M_{T} <100 GeV signal: M_{T} >100 GeV normalisation region: 100 GeV< $\not{\!\!E}_{T}$ <150 GeV





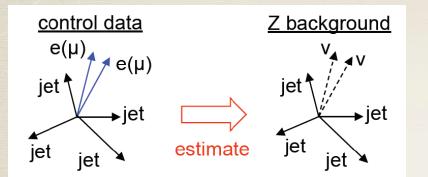
Epiphany 2010

The "tiles method" - yet another way Variables in the SM backgrounds may exhibit correlations SM shapes (fractions) must be known (eg from MC)



p18

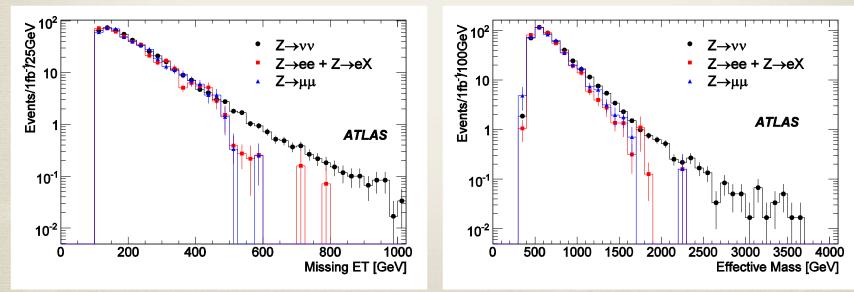
Z->vv + jets an important background to O-lepton > Replace method relies on the measured Z->I+I-

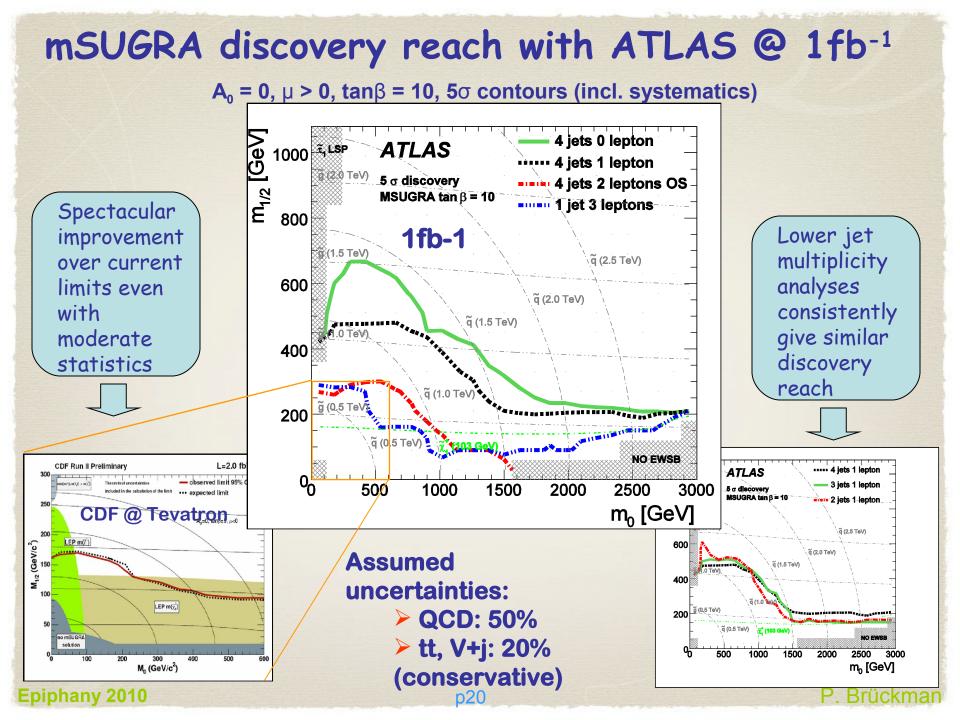


Epiphany 2010

Standard 0-lepton selection + Z->I⁺I⁻ with $p_{T}(I^{+}I^{-})$ substitution for $\not E_{T}$

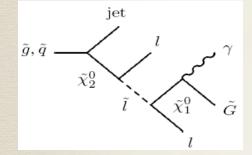
acceptance (η, p_T) , efficiency, and Br corrections must be applied!





mSUGRA is merely a convenient framework for assessing the discovery potential for R-conserving SUSY with χ^{0}_{1} as LSP. Other SUSY breaking scenarios lead to different EW-scale phenomenology. Will shortly discuss GMSB:

* with gravitino LSP and χ^{0}_{1} NLSP



name	NLO (LO) σ [pb]	Λ [TeV]	M_m [TeV]	C_G	$c\tau$ [mm]	$M_{\tilde{\chi}^0_1}$ [GeV]
GMSB1	7.8 (5.1)	90	500	1.0	1.1	118.8
GMSB2	7.8 (5.1)	90	500	30.0	$9.5 \cdot 10^{2}$	118.8
GMSB3	7.8 (5.1)	90	500	55.0	$3.2 \cdot 10^{3}$	118.8

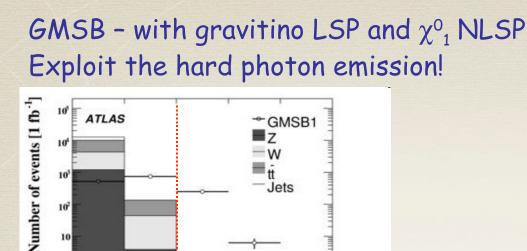
 $\chi_{\,1}^{\,0} o \,\gamma + ilde{G}$ lifetime

p21

* with gravitino LSP and meta-stable slepton NLSP

name	NLO (LO) σ [pb]	Λ [TeV]	M_m [TeV]	$M_{\tilde{\tau}_1}$ [GeV]
GMSB5	21.0 (15.5)	30	250	102.3

Others (not covered here): o Split SUSY (stable R-hadrons) o NUHM o AMSM

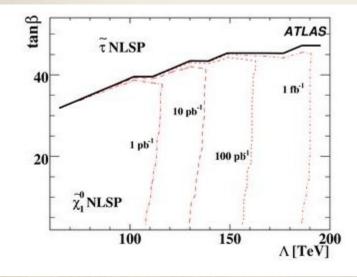


Discovery reach:

1

10

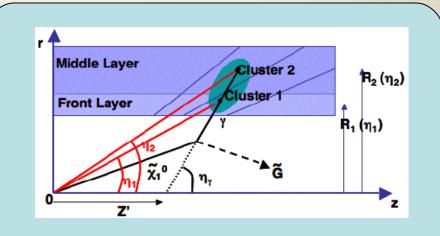
0



2

N.

 \geq 2 isolated photons p_T>20GeV <mark>≥4 jets p₊>50GeV</mark> ≥1 jet p₋>100GeV <mark>∉₇>100GeV</mark> ₽__>0.2M___ trigger: g55 || 2g17i



If χ^{0}_{1} lifetime long enough photons will appear as "non-pointing" Lifetime measurement possible from:

- Z' reconstruction
- Calorimeter timing

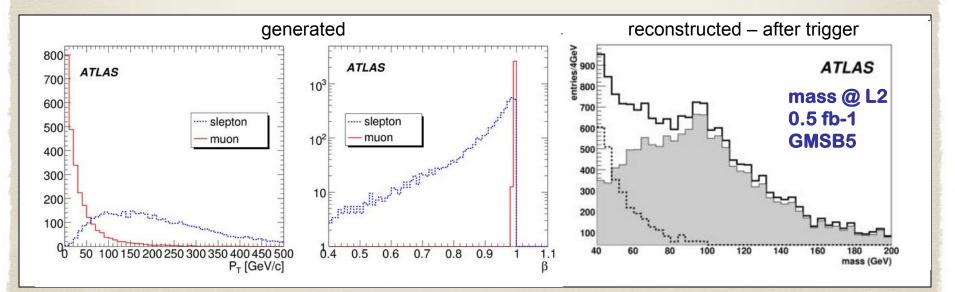
GMSB - with gravitino LSP and slepton NLSP Exploit the low velocity of the heavy meta-stable particle!

Signature: penetrating tracks with high p_{τ} and low β Signal in parts of detector in different bunch crossings: Online:

Epiphany 2010

L1: require regular muon high trigger (95% efficient) >L2: use 3ns resolution TOF information from RPC's (barrel only, ~50% eff.) Trigger on high mass! pT>40 GeV, β**<**0.97.

m>40 GeV



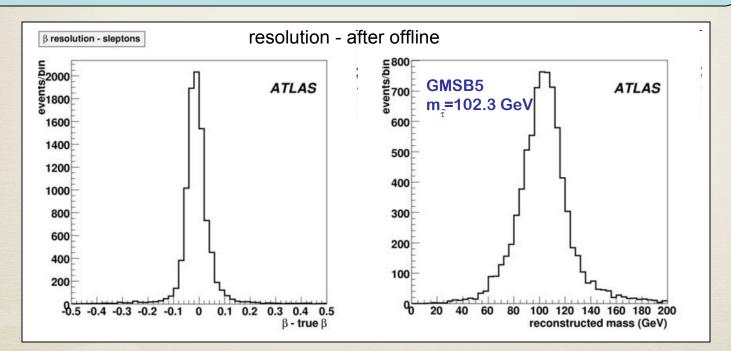
p23

GMSB - with gravitino LSP and slepton NLSP Exploit the low velocity of the heavy meta-stable particle!

The incorrect time of arrival distorts drift time measurements in the MDT's. Offline:

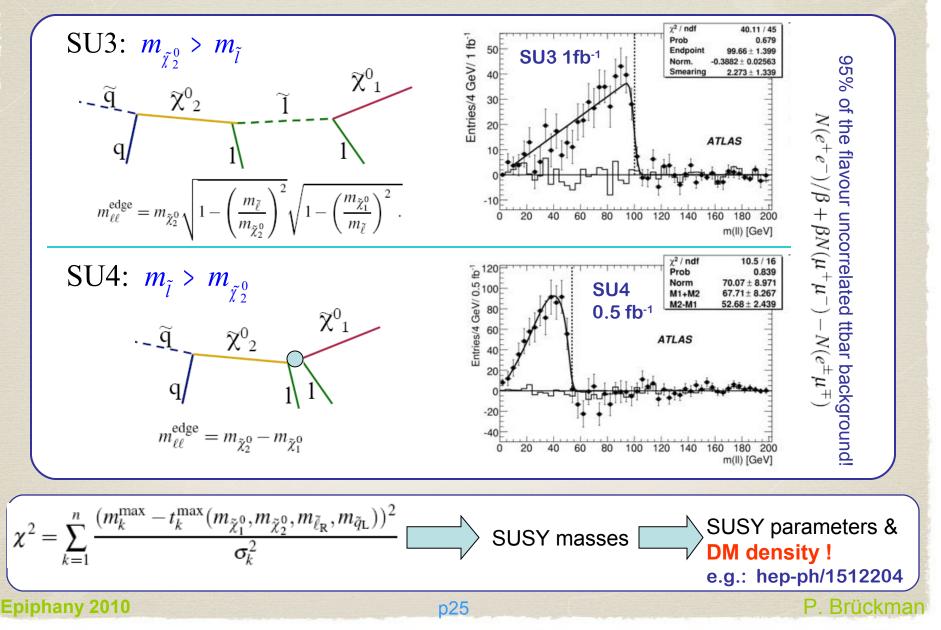
> Minimise the track reconstruction χ^2 w.r.t. tof.

>Combine this with L2 trigger to extract β .

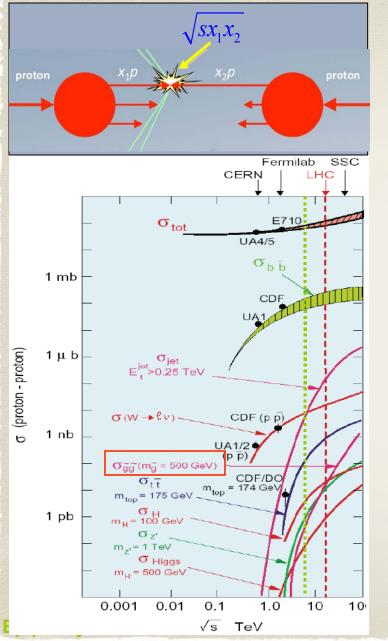


Note: similar technique applicable to R-hadron searches

Once SUSY discovered we will try to make a better acquaintance •sparticle masses, spin, couplings, etc.



What to expect this year? ~0.1fb⁻¹ @ 7(?)TeV



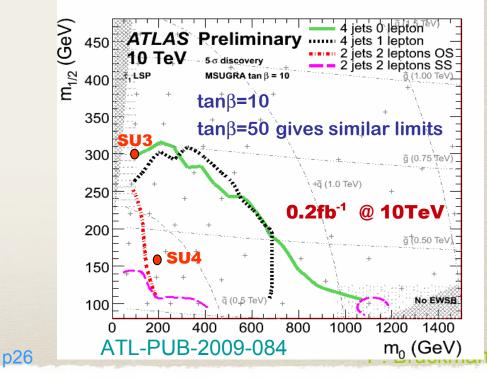
Cross-section steeply falling for heavy object production (e.g. sparticle pair) Predictions @ lower \sqrt{s} and lower L_{int} nontrivial:

o x-sections change rapidly

o background systematics are L_{int} dependent

If lucky we may see a glimpse of low mass SUSY (SU4-like)...

Recent update to 200fb⁻¹ @ 10TeV



Changing Gears



©We are all very excited about it ☺



LHC milestones - from Steve Myers

Date	Day	Achieved
Nov 20	1	Each beam circulating. Key beam instrumentation working.
Nov 23	4	First collisions at 450 GeV. First ramp (reached 560 GeV).
Nov 26	7	Magnetic cycling established (reproducibility).
Nov 27	8	Energy matching.
Nov 29	10	Ramp to 1.18 TeV.
Nov 30	11	Experiment solenoids on.
Dec 04	15	Aperture measurement campaign finished. LHCb and ALICE dipoles on.
Dec 05	16	Machine protection (Injection, Beam dump, Collimators) ready for safe operation with pilots.
Dec 06	17	First collisions with STABLE BEAMS, 4 on 4 pilots at 450 GeV, rates around 1Hz. Santa!
Dec 08	19	Ramp colliding bunches to 1.18 TeV
Dec 11	22	Collisions with STABLE BEAMS, 4 on 4 at 450 GeV, > 10 ¹⁰ per bunch, rates around 10Hz.
Dec 13	24	Ramp 2 bunches per beam to 1.18 TeV. Collisions for 90mins.
Dec 14	25	Collisions with STABLE BEAMS, 16 on 16 at 450 GeV, > 10 ¹⁰ per bunch, rates around 50Hz.
Dec 16	27	Ramp 4 on 4 to 1.18 TeV. Squeeze to 7 m.
Epiphany 2010		p29 P. Brückman

© 20-23 November in ATLAS ©

Friday, November 20:

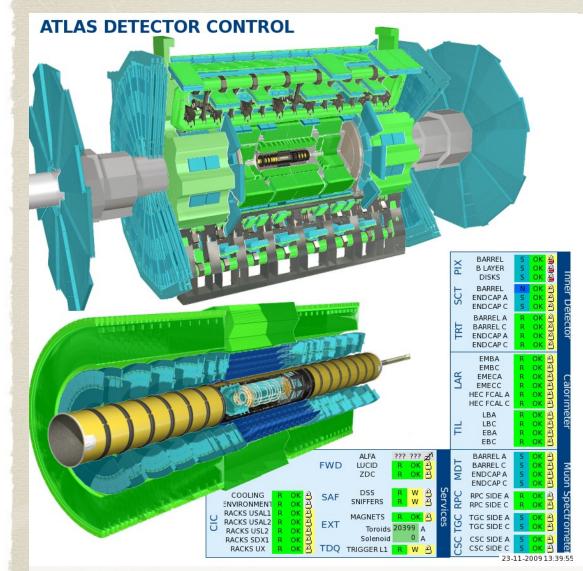
- ~ 20:30h: beam-1 threading \rightarrow 6 beam splashes to ATLAS
- ~ 22:30h: beam-2 threading \rightarrow 7 beam splashes to ATLAS

Saturday, 21 November:

- ~ 1h: beam-2 splashes to ATLAS \rightarrow 27 events (side C)
- ~ 4h: beam-1 splashes to ATLAS \rightarrow 26 events (side A)
- Sunday, 22 November:
- ~ 6h: 15 splash events to test beam abort by BCM \rightarrow successful
- Monday 23 November:
- ~ 6:30: last series of splashes to ATLAS \rightarrow 25 events (side C)
- ~ 13:30: two beams injected for collisions at IP1 and IP5
- ~ 14:22: first ATLAS collision event seen !!!

P. Brückm

Detector fully operational

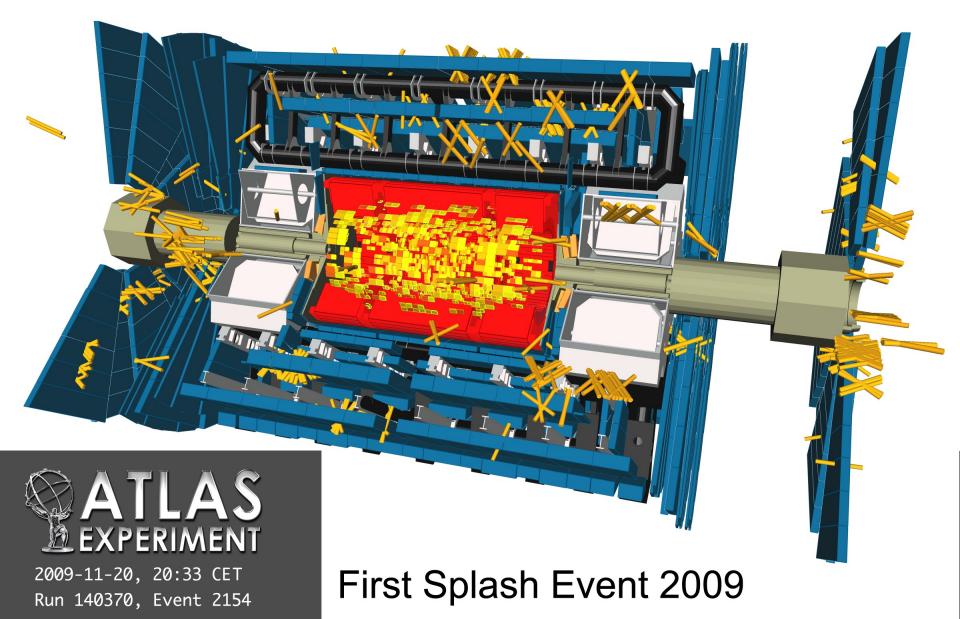


p31

Setup for beam commissioning

- Pixel off
- SCT(standard bias voltage 150 V)
 - Standby V is 20 V →~50%
 hit efficiency (increases with incidence angle)
 - Barrel and endcap increased to 50V for short stable beam periods during collisions
 - Barrel voltage sometimes lower than 20V for beam set up (eg. splash events)
- All other systems ON
- No solenoid field, toroids ON

Beam One (clockwise)



1st Beam Splash from Beam-2 (anticlockwise)



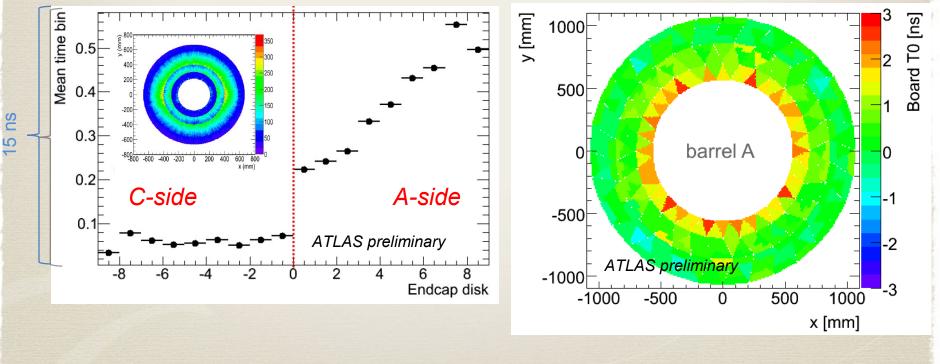
Timing studies with beam-splash events (Inner Detector)

Inner tracking systems:

Epiphany 2010

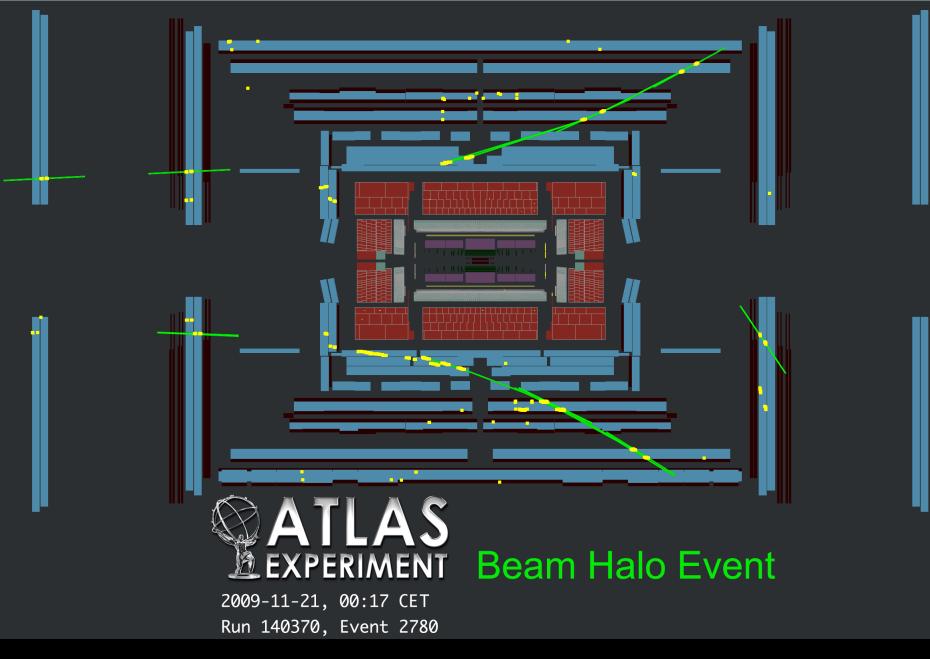
- SCT already well timed-in from cosmics and known cable lengths (better than 2 ns)
- TRT boards timed-in to better than 2 ns

Beam-1 arriving from A-side: timing as collisions for C-side, but wrong for A-side



p34

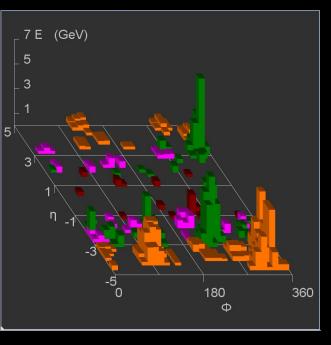
TRT Barrel: plot made with collision timing \rightarrow sensitive to ToF effect on Inner Boards !



A projection of a "Beam Halo" event showing tracks in the Muon Spectrometer.

Candidate Collision Event

J





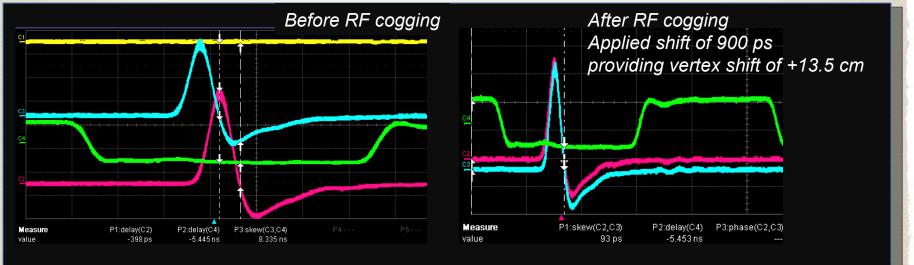
2009-11-23, 14:22 CET Run 140541, Event 171897

http://atlas.web.cern.ch/Atlas/public/EVTDISPLAY/events.html

10100000

Understanding accelerator operation - feedback to LHC

- The ATLAS beam pickups showed a phase inconsistency of 900 ps causing the primary vertex to be shifted by –13.5 cm in z
- Based on this information, at around 14:50, the LHC operators performed an RF cogging to correct the z positioning of the beam spot at IP1



Beam pickup scope shots, beam 1 & 2

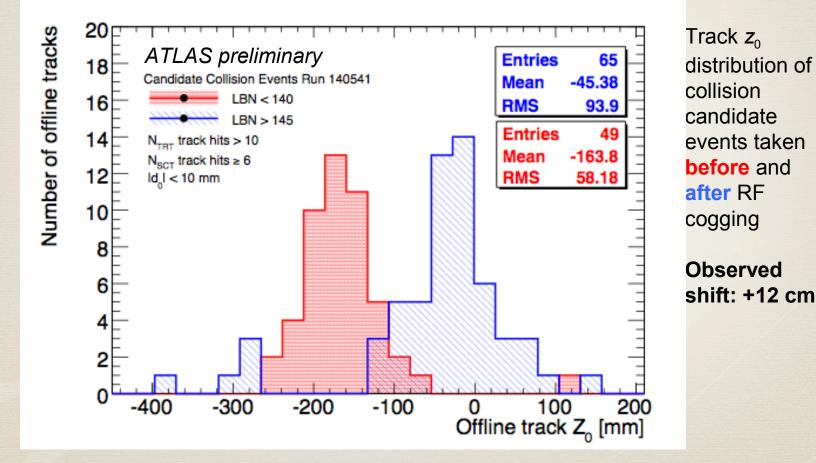
Bunches stable within 20 ps (RMS) !

P. Brückman

Epiphany 2010

Understanding accelerator operation - offline analysis

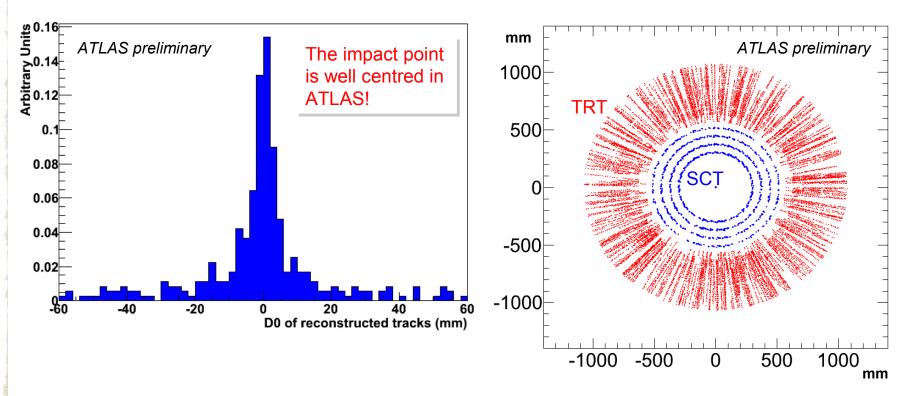
- ATLAS has taken data before and after the RF cogging
- Must observe shift in z₀of tracks if indeed we select collision events!



Very early days

Tracking(challenging w/o Pixel, limited SCT and solenoid field off!)

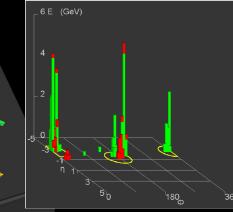
- Without solenoid field no separation of tracks by momenta
- Fit impact parameter in a "silver-plated" sample with SCT >= 20 V and number of SCT hits >= 6 (46 events)



Scatter plot of hits on tracks (barrel, 46 events)

A di-jet candidate

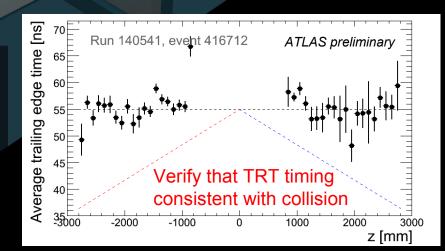
II INTANIA INTRA (KATRA)

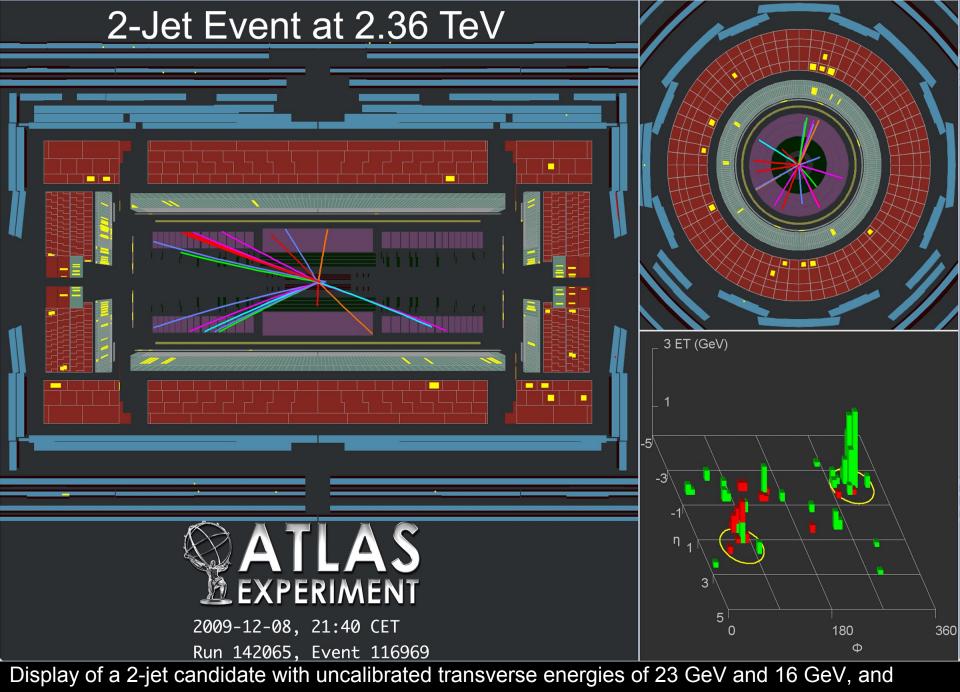


Run 140541 Event 416712

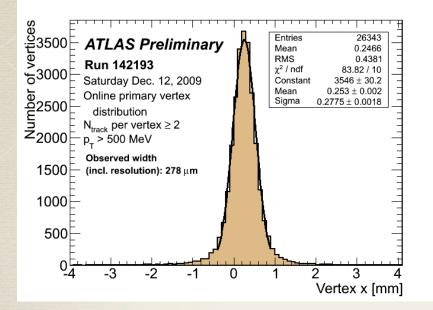
Two jets back-to-back in ϕ , both with (uncalibrated) $E_{\tau} \sim 10$ GeV, η of -1.3 and -2.5, \sim no missing E_{τ}

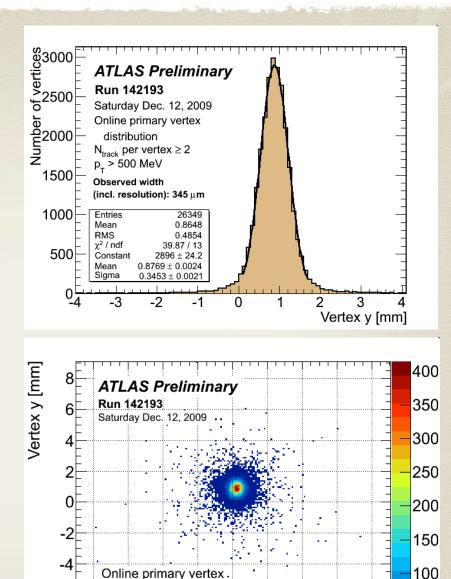
Triggered by MBTS A/B in time, several hits Also triggered by L1Calo EM3





pseudo-rapidities of -2.1 and 1.4, respectively.





distribution

p_ > 500 MeV

 N_{track} per vertex ≥ 2

-2

0

2

6

Vertex x [mm]

4

8

-6

-8

Beam position is monitored OnLine within the L2 HLT algorithms. BS size ~250 µm

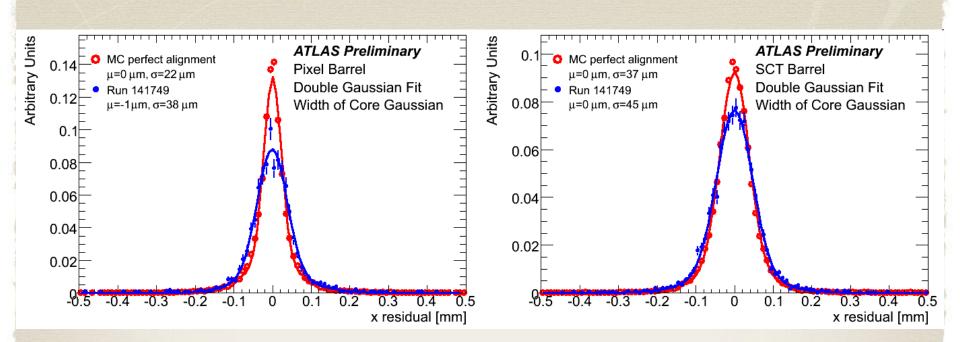
Epiphany 2010

p42

P. Brückman

50

n

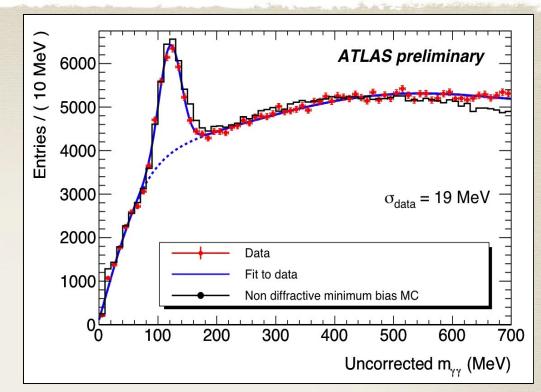


Silicon alignment obtained from analysis of cosmic events works surprisingly well for the collision tracks.

Note: cosmic muon illumination was very poor in the horizontal direction!

□ First attempts to realign the entire Inner Detector with collision data are now underway. End-cap disk alignment has already marked a spectacular improvement.

$\pi^0 - \gamma\gamma$



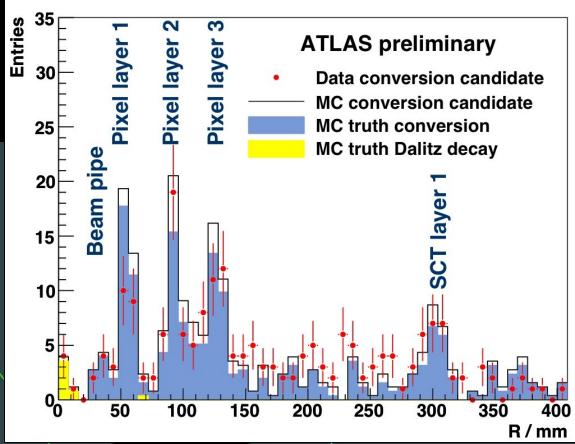
The analysis is based on EM topocluster (with 4 σ above noise for the seed, 3 σ for the seed neighbour and 0 σ for the next neighbour). π° selection cuts :

- E, of each cluster > 300 MeV
- E_{t} of π^{0} candidate > 900 MeV
- Shower shapes compatible with photons

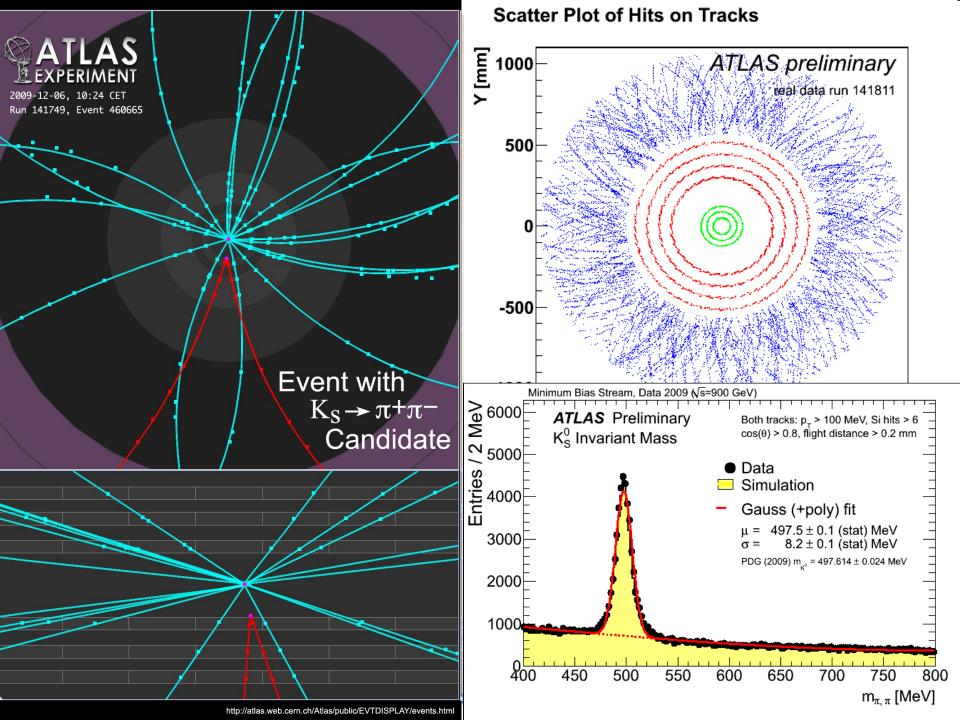
All combinatorial pairing are included in the distribution. The distribution is normalized to the number of entries. The cells are only calibrated at the electromagnetic energy scale (no dead material correction applied).

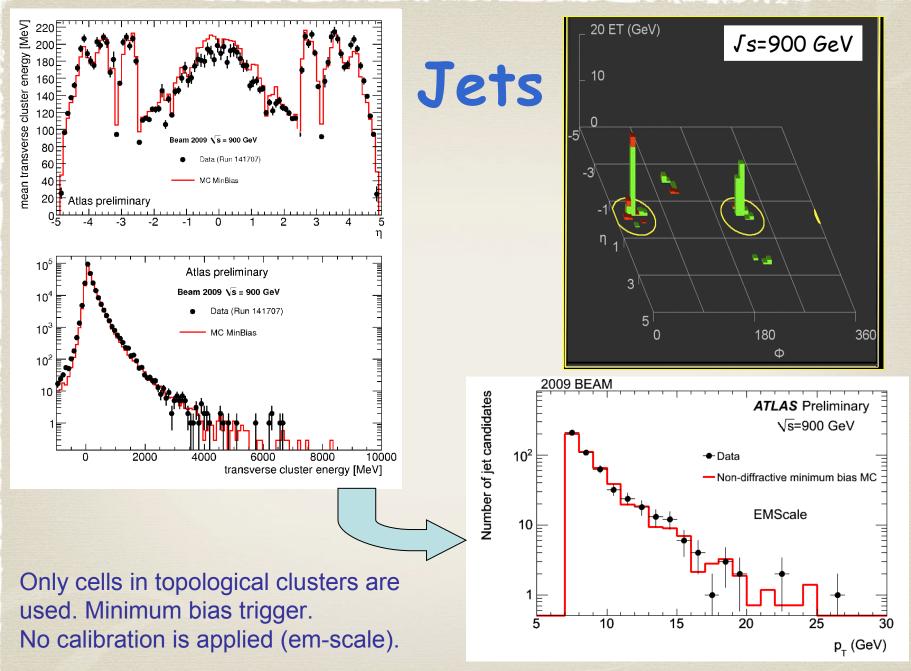
Epiphany 2010









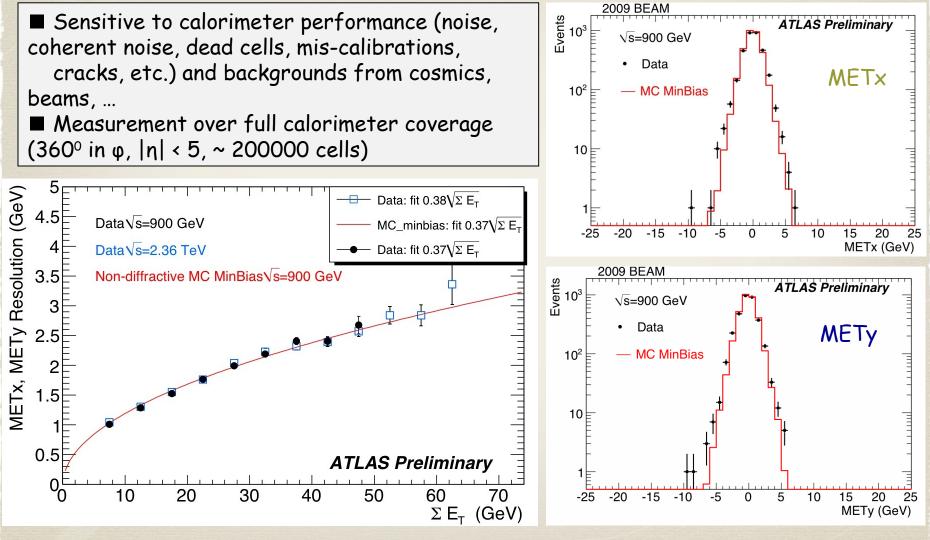


p47

Epiphany 2010

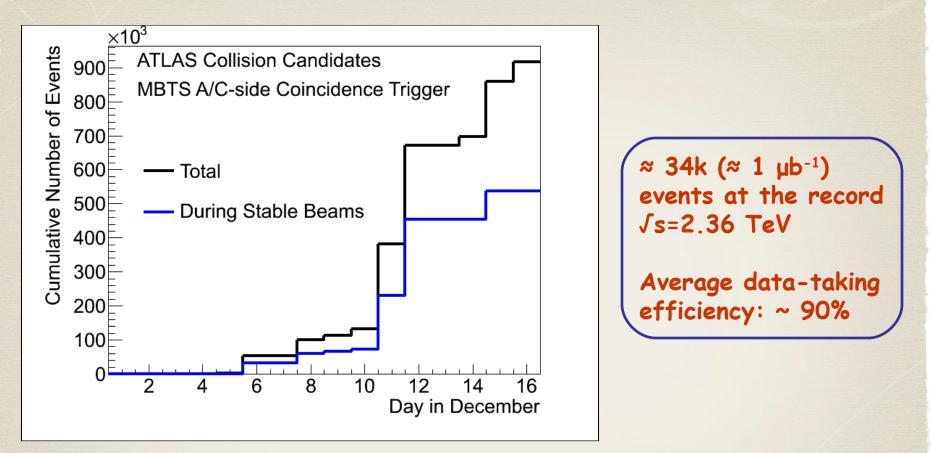
Missing transverse energy (essential for SUSY searches!)

Epiphany 2010



p48

Statistics of collision events collected to date



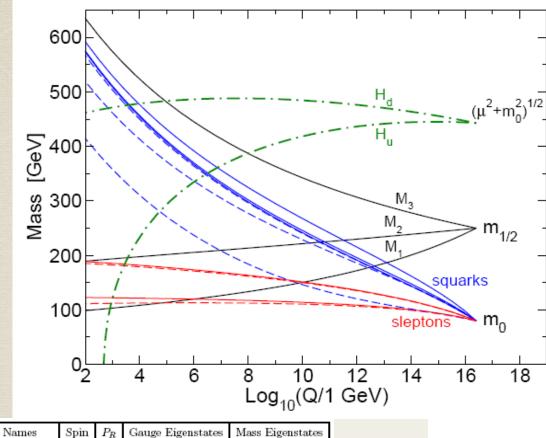
LHC stopped for the Christmas shut-down on 16/12
 Restart planned for mid-February
 2010 goal: 3.5TeV on 3.5TeV (+?), Lumi<10³², intL~100pb⁻¹

Epiphany 2010

Thank you



BACKUP SLIDES



RG evolution of scalar and gaugino masses in a typical SUGRA

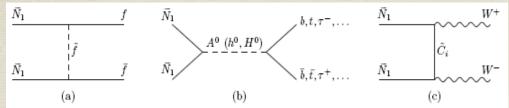
Typical sparticle mass spectrum of SUGRA

Mass

Eigenstates					$\overline{\tilde{g}}$	$\tilde{d}_L \tilde{u}_L$	\tilde{t}_2	
$I^0 A^0 H^{\pm}$					g		\tilde{b}_2	
(same)		H^{\pm}	\tilde{N}_4			$\tilde{u}_R \tilde{d}_R$	~	
(same)		$\overline{H^0 A^0}$	$\frac{\tilde{N}_4}{\tilde{N}_3}$	\tilde{C}_2			b_1	
$\overline{t}_2 \ \overline{b}_1 \ \overline{b}_2$			<i>I</i> V3	\mathbb{C}_2			ĩ	
(same)							\tilde{t}_1	
same)								
$\overline{\tau}_2 \overline{\nu}_{\tau}$						\tilde{e}_L	$ ilde{ au}_2$	
$\overline{N}_2 \ \overline{N}_3 \ \overline{N}_4$		h^0	\tilde{N}_2	\tilde{C}_1		$\tilde{\nu}_e$	$\tilde{\nu}_{\tau}$	
$\vec{c}_{1}^{\pm} \vec{C}_{2}^{\pm}$						\tilde{e}_R	$\overline{\tilde{ au}_1}$	
(same)			\tilde{N}_1				· 1	
same)	05							

1.0	rames	opm	1 K	Gadge Eigenstates	Mass Eigenstates	
-	Higgs bosons	0	+1	$H^0_u \ H^0_d \ H^+_u \ H^d$	h^0 H^0 A^0 H^{\pm}	
-		0	-1	$\vec{u}_L \ \vec{u}_R \ \vec{d}_L \ \vec{d}_R$	(same)	
Whether war and an an an an an	squarks			$\vec{s}_L \ \vec{s}_R \ \vec{c}_L \ \vec{c}_R$	(same)	
				$\overline{t}_L \ \overline{t}_R \ \overline{b}_L \ \overline{b}_R$	$\overline{t}_1 \ \overline{t}_2 \ \overline{b}_1 \ \overline{b}_2$	
		0	-1	$\overline{e}_L \ \overline{e}_R \ \overline{\nu}_e$	(same)	
	sleptons			$\bar{\mu}_L \ \bar{\mu}_R \ \bar{\nu}_\mu$	(same)	
				$\bar{\tau}_L \ \bar{\tau}_R \ \bar{\nu}_\tau$	$\vec{\tau}_1 \ \vec{\tau}_2 \ \vec{\nu_\tau}$	
	neutralinos	1/2	-1	$\overline{B}^0 \ \widetilde{W}^0 \ \overline{H}^0_u \ \overline{H}^0_d$	$\bar{N_1}\ \bar{N_2}\ \bar{N_3}\ \bar{N_4}$	
	charginos	1/2	-1	\widetilde{W}^{\pm} \overline{H}^{+}_{u} \overline{H}^{-}_{d}	\bar{C}_{1}^{\pm} \bar{C}_{2}^{\pm}	
	gluino	1/2	-1	\overline{g}	(same)	
Perfor	goldstino (gravitino)	1/2 (3/2)	-1	\tilde{G}	(same)	

LSP (NLSP) annihilation mechanisms:



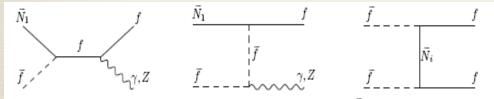
a) bulk regionb) funnel regionc) focus point

Figure 9.13: Contributions to the annihilation cross-section for neutralino dark matter LSPs from (a) t-channel slepton and squark exchange, (b) near-resonant annihilation through a Higgs boson (s-wave for A^0 , and p-wave for h^0 , H^0), and (c) t-channel chargino exchange.



focus point

Figure 9.14: Some contributions to the co-annihilation of dark matter \bar{N}_1 LSPs with slightly heavier \bar{N}_2 and \bar{C}_1 . All three diagrams are particularly important if the LSP is higgsino-like, and the last two diagrams are important if the LSP is wino-like.



sfermion co-annihilation

P. Brückn

Figure 9.15: Some contributions to the co-annihilation of dark matter N_1 LSPs with slightly heavier sfermions, which in popular models are most plausibly staus (or perhaps top squarks).

Epiphany 2010

Simulation of SUSY signal and backgrounds

- ttbar: MC@NLO + HERWIG for fragmentation
- W+jets, Z+jets: ALPGEN + HERWIG + NNLO k-factors
- N<4 jets PYTHIA for W & Z backgrounds</p>
- WW, ZZ, WZ: HERWIG + NLO k-factors
- SUSY benchmark points: PROSPINO 2.0.6 + CTEQ6M
- SUSY scans: ISAJET 7.75

SUSY scan strategy

- Scan of the parameter space using fast detector simulation (ATLFAST)
- Only direct exclusions from LEP&Tevatron are considered (not cosmological!)
- Pile-up not included
- Signal uses LO HERWIG without further corrections (CONSERVATIVE!)
- Background from the full simulation + assumption about the DD methods (50% QCD, 20% W, Z & top) – NLO Monte Carlo
- Separate grids for mSUGRA tan β =10, tan β =50, GMSB, NUHM...
- "Multiple comparisons" correction to significance is always applied.
- Significance (Z_n) is given by the convolution of Poisson distribution

(statistical fluctuation) and a Gaussian (systematic uncertainty):

$$Z_n = \sqrt{2} \operatorname{erf}^{-1}(1-2p) \qquad \text{with} \quad p = A \int_0^\infty db \ G(b; N_b, \delta N_b) \sum_{i=N_{\text{data}}}^\infty \frac{e^{-b}b^i}{i!},$$

ποιπαπρατιστ

