



***Fitting PDFs with  
Monte Carlo solutions of evolution equations  
in QCD***

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New parton shower MC project is currently constructed by IFJ-PAN group from Krakow, based on evolution in rapidity space

- Such parton shower with non-standard evolution (CCFM-like,  $\alpha(Q(1-z))$ ,  $\alpha(Q(x\frac{1-z}{z}))$ ) etc) requires dedicated PDF fits as a part of the project
- MC is slow by its nature (permille accuracy  $\approx 3$  days of running). Therefore MC cannot be used directly for fitting - fast procedure must be designed

# Evolution equation

$$\partial_t D_A(x, t) = P_{AB}(x, t) \otimes^x D_B(x, t), \quad (1)$$

$$D_A(x, t = 0) = D_A^0(x, \alpha_1^A, \dots, \alpha_k^A). \quad (2)$$

$D_A(x, t)$  denotes PDF of the type  $A$  with  $A = q_i, \bar{q}_i, g; i = 1, \dots, n$ . and  $t$  is the evolution time. The splitting kernels  $P_{AB}(x, t)$  include also coupling constants and the convolution symbol  $\otimes^x$  stands for

$$(f(x, \alpha) \otimes^x g(x, \beta))(x) = \int_x^1 dy dz \delta(x - yz) f(y, \alpha) g(z, \beta). \quad (3)$$

# Decomposition

The decomposition of the solution  $D^A(x, t)$  can be written as follows

$$D_A(x, t) = D_{AB}^\delta(x, t) \otimes^x D_B^0(x, \alpha_1^B, \dots, \alpha_k^B) \quad (4)$$

$$\partial_t D_{AB}^\delta(x, t) = P_{AC}(x, t) \otimes^x D_{CB}^\delta(x, t), \quad (5)$$

$$D_{AB}^\delta(x, 0) = \delta_{AB} \delta(1 - x). \quad (6)$$

- Whole dependence of fitted parameters ( $\alpha_j^B$ ) is limited to  $D_B^0$
- Only single MC run needed for  $D_{AB}^\delta$  (independent of  $\alpha_j^B$ )

# *Integration procedure*

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- Convolution  $\otimes^x$  is a 1-dim integral, can be done numerically in a fast and accurate way.
- MC histogram of  $D^\delta$  is parametrized by polynomial second order functions.
- Because we use log-scale for  $x$ , value at  $x$  close to 1 must be extrapolated from neighbouring bins
- This extrapolation could be avoided if linear scale were used, but it is not necessary for  $10^{-3}$  precision

# *Test of integration procedure*

We use 2-dim system of gluon (G) and singlet ( $\Sigma$ ) PDF-s

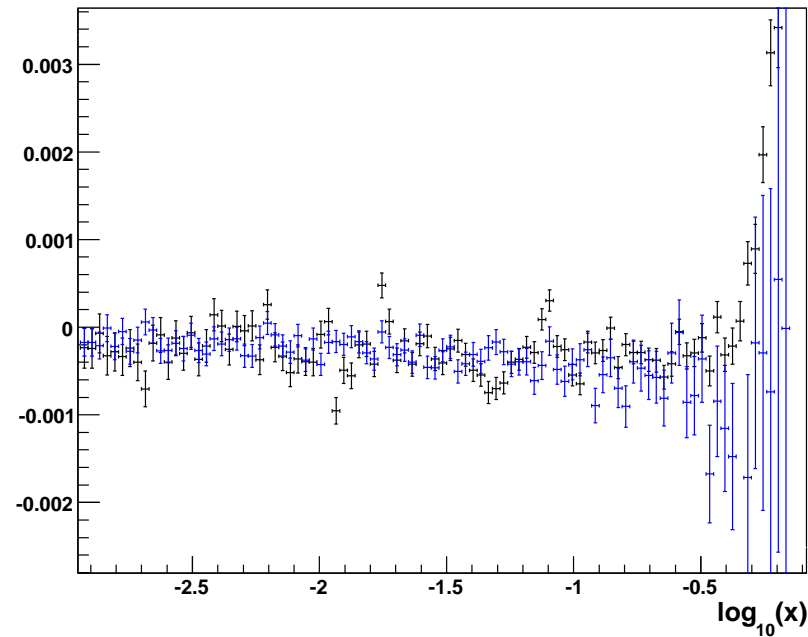
$$\begin{aligned} D_G(x, t) &= D_{GG}^\delta(x, t) \otimes D_G^0(x) + D_{G\Sigma}^\delta(x, t) \otimes D_\Sigma^0(x), \\ D_\Sigma(x, t) &= D_{\Sigma G}^\delta(x, t) \otimes D_G^0(x) + D_{\Sigma\Sigma}^\delta(x, t) \otimes D_\Sigma^0(x). \end{aligned} \tag{7}$$

Initial conditions

$$\begin{aligned} D_G^0(x) &= 1.908 \cdot x^{-1.2} (1-x)^{5.0}, \\ D_\Sigma^0(x) &= D_{sea}^0(x) + D_u^0(x) + D_d^0(x), \\ D_{sea}^0(x) &= 0.6733 \cdot x^{-1.2} (1-x)^{7.0}, \\ D_u^0(x) &= 2.187 \cdot x^{-0.5} (1-x)^{3.0}, \\ D_d^0(x) &= 1.230 \cdot x^{-0.5} (1-x)^{4.0}. \end{aligned} \tag{8}$$

# *Test of convolution Gluon and Singlet (LO)*

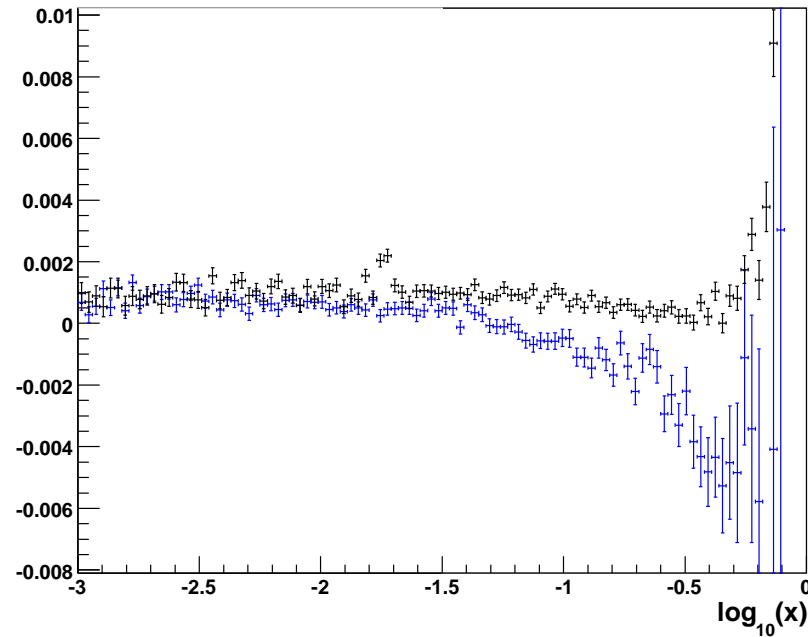
Ratio of gluon (blue) and singlet (black) generated directly from initial distribution and obtained from convolution of  $D_\delta$ . Evolution from  $Q = 1$  to  $Q = 100\text{GeV}$ , LO type



Agreement  $5 \cdot 10^{-4}$  except of  $x \sim 1$ , where  $D$  is very small any way

# Test of convolution Gluon (NLO)

Ratio of gluon (blue) and singlet (black) generated directly from initial distribution and obtained from convolution of  $D_\delta$ . Evolution from  $Q = 1$  to  $Q = 100\text{GeV}$ , NLO type



Agreement  $1 \times 10^{-3}$  except of  $x \sim 1$ , where  $D$  is very small anyway



- To check correctness of the procedure we fitted PDFs to PDFs
- We used the same  $G - \Sigma$  system
- We used MINUIT and minimized  $\chi^2$

$$\chi^2(\alpha_1^G, \dots, \alpha_k^G; \alpha_1^\Sigma, \dots, \alpha_k^\Sigma) = \sum_{A=G, \Sigma} \sum_i \frac{(D_A^X(x_i, t, \alpha_1^A, \dots, \alpha_k^A) - D_A^Y(x_i, t))^2}{e_A^Y(x_i, t)^2} \quad (9)$$

X and Y denote evolution type: if  $X = Y$  - technical test, if  $X \neq Y$  - physical difference

# Fitting PDF to PDF-test (LO)

	$\alpha_1^G$	$\alpha_2^G$	$\alpha_3^G$	$\alpha_1^u$	$\alpha_2^u$	$\alpha_3^u$	$\alpha_1^d$	$\alpha_2^d$	$\alpha_3^d$
I	1.908	1.200	5.000	2.187	0.500	3.000	1.230	0.500	4.000
F	1.907	1.199	4.994	2.186	0.501	3.008	1.229	0.505	3.991

	$\alpha_1^{sea}$	$\alpha_2^{sea}$	$\alpha_3^{sea}$
I	0.67	1.200	7.00
F	0.68	1.199	7.06

Comparison of fitted values of coefficients with the original ones used for generation for the LO evolution at  $Q = 100$  GeV. I – input values, F – fitted values. The initial distributions are  $\alpha_1^A x^{-\alpha_2^A} (1-x)^{\alpha_3^A}$ . Accuracy mostly within  $2\sigma$ .

# Fitting PDF to PDF-test (NLO)

	$\alpha_1^G$	$\alpha_2^G$	$\alpha_3^G$	$\alpha_1^u$	$\alpha_2^u$	$\alpha_3^u$	$\alpha_1^d$	$\alpha_2^d$	$\alpha_3^d$
I	1.908	1.200	5.000	2.187	0.500	3.000	1.230	0.500	4.000
F	1.905	1.200	5.002	2.187	0.503	3.013	1.230	0.513	3.997

	$\alpha_1^{sea}$	$\alpha_2^{sea}$	$\alpha_3^{sea}$
I	0.673	1.200	7.000
F	0.678	1.198	7.094

Comparison of fitted values of coefficients with the original ones used for generation for the NLO evolution at  $Q = 100$  GeV. I – input values, F – fitted values. The initial distributions are  $\alpha_1^A x^{-\alpha_2^A} (1-x)^{\alpha_3^A}$ . Accuracy mostly within  $2\sigma$ .

# Fitting PDF to PDF

Fitting different evolutions: Generated LO DGLAP with  $\alpha(Q(1 - z))$ , fitted by NLO DGLAP

	$\alpha_1^G$	$\alpha_2^G$	$\alpha_3^G$	$\alpha_1^u$	$\alpha_2^u$	$\alpha_3^u$	$\alpha_1^d$	$\alpha_2^d$	$\alpha_3^d$
I	1.908	1.200	5.000	2.187	0.500	3.000	1.230	0.500	4.000
F	1.801	1.202	4.724	1.992	0.359	2.625	0.46	0.234	2.991

	$\alpha_1^{sea}$	$\alpha_2^{sea}$	$\alpha_3^{sea}$
I	0.673	1.200	7.000
F	0.605	1.47	8.011

$Q = 100$  GeV. I – input values, F – fitted values. The initial distributions are  $\alpha_1^A x^{-\alpha_2^A} (1 - x)^{\alpha_3^A}$  The change in parameters is big (up to 50%)

In LO we have

$$F_2(x, t) = \sum_{A=q, \bar{q}} \int_0^1 d\xi D_A(\xi, t) x e_A^2 \delta(x - \xi) = \sum_{A=q, \bar{q}} D_A(x, t) x e_A^2. \quad (10)$$

$$F_2(x, t, \vec{\alpha}_1, \dots, \vec{\alpha}_k) = x \sum_{A=q, \bar{q}} e_A^2 D_{AB}^\delta(x, t) \otimes^x D_B^0(x, \alpha_1^B, \dots, \alpha_k^B). \quad (11)$$

and we fit with MINUIT with  $\chi^2$

$$\chi^2(a, b, \dots) = \sum_n \sum_i \frac{(F_2(x_i, t_n, \vec{\alpha}_1, \dots, \vec{\alpha}_k) - F_{2exp}(x_i, t_n))^2}{e_{F_{2exp}}(x_i, t_n)^2}. \quad (12)$$

# Fitting F2

In NLO nontrivial coefficient functions appear

$$F_2^{NLO}(x, t) = x \sum_{A=q, \bar{q}} e_A^2 D_A(x, t) + \Delta F_2, \quad (13)$$

$$\Delta F_2 = x \sum_{A=q, \bar{q}} e_A^2 D_A(x, t) \otimes^x \frac{\alpha_s}{2\pi} C_A^{\overline{MS}}(x) + x \sum_{A=q, \bar{q}} e_A^2 D_g(x, t) \otimes^x \frac{\alpha_S}{2\pi} C_g^{\overline{MS}}(x), \quad (14)$$

where

$$C_q^{\overline{MS}} = C_F \left( 2 \left( \frac{\ln(1-z)}{1-z} \right)_+ - \frac{3}{2} \left( \frac{1}{1-z} \right)_+ - (1+z) \ln(1-z) - \frac{1+z^2}{1-z} \ln(z) \right. \\ \left. + 3 + 2z - \left( \frac{\pi^2}{3} + \frac{9}{2} \right) \delta(1-z) \right) \quad (15)$$

$$C_g^{\overline{MS}} = T_R \left( ((1-z)^2 + z^2) \ln \left( \frac{1-z}{z} \right) - 8z^2 + 8z - 1 \right). \quad (16)$$

# Fitting $F_2$ to $F_2$ obtained from QCDNum16 - test (LO)

Fitting two structure functions:  $F_2$  (LO) obtained from Monte Carlo to  $F_2$  from QCDNum16

	$\alpha_1^G$	$\alpha_2^G$	$\alpha_3^G$	$\alpha_1^u$	$\alpha_2^u$	$\alpha_3^u$	$\alpha_1^d$	$\alpha_2^d$	$\alpha_3^d$
I	1.908	1.200	5.000	2.187	0.500	3.000	1.230	0.500	4.000
F	1.908	1.199	4.988	2.187	0.502	3.07	1.241	0.507	4.08

	$\alpha_1^{sea}$	$\alpha_2^{sea}$	$\alpha_3^{sea}$
I	0.673	1.200	7.000
F	0.674	1.200	6.944

$Q = 100$  GeV. I – input values, F – fitted values. The initial distributions are  $\alpha_1^A x^{-\alpha_2^A} (1-x)^{\alpha_3^A}$

# Fitting $F_2$ to $F_2$ obtained from QCDNum16 - test (NLO)

Fitting two structure functions:  $F_2$  (NLO) obtained from Monte Carlo to  $F_2$  from QCDNum16

	$\alpha_1^G$	$\alpha_2^G$	$\alpha_3^G$	$\alpha_1^u$	$\alpha_2^u$	$\alpha_3^u$	$\alpha_1^d$	$\alpha_2^d$	$\alpha_3^d$
I	1.908	1.200	5.000	2.187	0.500	3.000	1.230	0.500	4.000
F	1.908	1.199	4.979	2.196	3.06	0.503	1.25	0.51	4.04

	$\alpha_1^{sea}$	$\alpha_2^{sea}$	$\alpha_3^{sea}$
I	0.673	1.200	7.000
F	0.675	1.200	6.86

$Q = 100$  GeV. I – input values, F – fitted values. The initial distributions are  $\alpha_1^A x^{-\alpha_2^A} (1-x)^{\alpha_3^A}$



- We have developed framework for fitting PDFs from Monte Carlo solutions of evolution equations.
- It is based on factorization property of the evolution equations.
- It is very fast and numerically precise.
- Numerical tests have been successfully performed at the level of PDFs and at the level of F2.