

# CRs in the 'Knee'-Region - Some Results from KASCADE -

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Astrophysics Motivation: Quest of the Knee **Experimental** Selected Results on: Tests of high-energy interaction models > CR energy spectra & Composition > sources of systematic uncertainties Brief Status of KASCADE-Grande Perspectives

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# Origin of the "Knee" ??

Maximum energy of accelerator ?



# Origin of the "Knee" ?? 2 Leakage from Galaxy ?



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Nucl. Instr. Meth. A513 (2003) 490

#### (Karlsruhe Shower Core and Array Detector)





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### **KASCADE** Central Detector



J. Engler at al., NIM A427 (1999) 528

# Myon-Tracking Detector

#### $(50 \times 5.5 \text{ m}^2)$



# Single Airshower in $e/\gamma$ -Detectors





x (m)

0

104

y(m)

0

-104

# Single Events in Calorimeter...

#### Single Hadron (21 TeV)



# **Observables of KASCADE per shower**

- from Detector Array
  - shower direction  $(\theta, \phi)$ ,
  - shower core (x,y)
  - electron size (N\_)
  - muon size  $(N_{\mu}^{*})$
  - lateral particle distr. (s, R<sub>m</sub>)
- from µ tracking Detector
  - $\mu$  size > 800 MeV (N<sup>MTD</sup>)
  - local μ-density (ρ<sub>μ</sub>)
  - lateral µ-distr. (s<sub>µ</sub>, R<sub>µ</sub>)
  - $\mu$ -production height (h<sub>µ</sub>)

- <u>from Calorimeter</u>
  - number of hadrons with
    E > 100 GeV (N<sub>h</sub>)
  - sum of hadronic energy  $(\Sigma E_h)$
  - energy of leading hadron (Ehmax)
  - parameters of spatial hadron distribution ( $R_h$ ,  $\lambda$ , ...)
- from MWPC system
  - # of  $\mu s > 2.2 \text{ GeV}(N_{\mu}^{*})$
  - local μ-density (<mark>ρ</mark>μ)
  - hit pattern, 'spottiness'
- <u>from trigger plane</u>
  - # of μs > 490 MeV (Ν<sub>μ</sub><sup>ι</sup>)
  - local μ-density (ρ<sub>μ</sub>)
  - particle arrival times  $(\tau_{\mu})$

# **Observables of KASCADE per shower**





binning in N<sub>e</sub> → p-simulations should resemble data

binning in  $N_{\mu} \rightarrow$ 

*p/Fe simul. should be above/below data* 





binning in N<sub>e</sub> → p-simulations should resemble data

binning in N<sub>µ</sub> → p/Fe simul. should be above/below data



(N<sub>e</sub>, N<sub>µ</sub>) ⇔ (Energy, Mass)



#### **CORSIKA Simulations**

Data

 $(N_e, N_u) \Leftrightarrow (Energy, Mass)$ 



#### H. Ulrich, Doct. Thesis 2003

# **Test of Unfolding**



#### Synthetic input spectra generated by assuming 5 diff. primaries (p, He, C, Si, Fe) and a rigidity dependent knee



# Test of Unfolding



#### Synthetic input spectra generated by assuming 5 diff. primaries (p, He, C, Si, Fe) and a rigidity dependent knee



# All-Particle Spectrum



# E-Spectra of Mass-Groups



H. Ulrich, Doct. Thesis 2003

# E-Spectra of Mass-Groups



# **QGSJet vs SIBYLL**





# **QGSJet**

Schaueranzahl







u-deficit a/o electron excess in SIBYLL



### Direct Measurements vs KASCADE-Data



# KASCADE vs JACEE & RUNJOB



### Mass from Muon Production Height





U



C. Büttner, Doct. Thesis 2003

Light → Heavy



### **Conclusions so far:**

- Unfolding Method(s) work properly
- Spectra of light elements show a distinct knee at ~ 2 PeV
- Rigidity dependent knee is favoured, but...

Support for astrophysical interpretation of knee

Interaction Models still not satisfactory

• "Optimal" Model to describe KASCADE-Data:

- Lower  $N_{\mu}$  at PeV energies
- more rapid increase of  $N_{\mu}$  towards higher energies *likewise, but less favoured:*
- Higher N at PeV energies
- 🤄 clues to better hadronic interaction models

# Towards higher Energies: KASCADE-Grande



Example of Joint Event

251 KASCADE 37 Grande Stations

time

arrival

 $E > 10^{17} \text{ eV}$ 

µ-detector, run 004358 event 0160542

# **Combined Lateral Distribution**



### **Radio-Observations of EAS**

Long standing question: Are EAS observable by their radio signal ? (30-80 MHz) observe EAS at their maximum, 24 hrs a day!

**KASCADE-Grande used as a reference and trigger** 



<u>Status:</u> 10 antennes in field, triggered by KASCADE



in Collab. with H. Falcke et al.





#### KASCADE-Grande Collaboration

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